

# $\nu_\mu$ Disappearance

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“Let’s talk about it!”

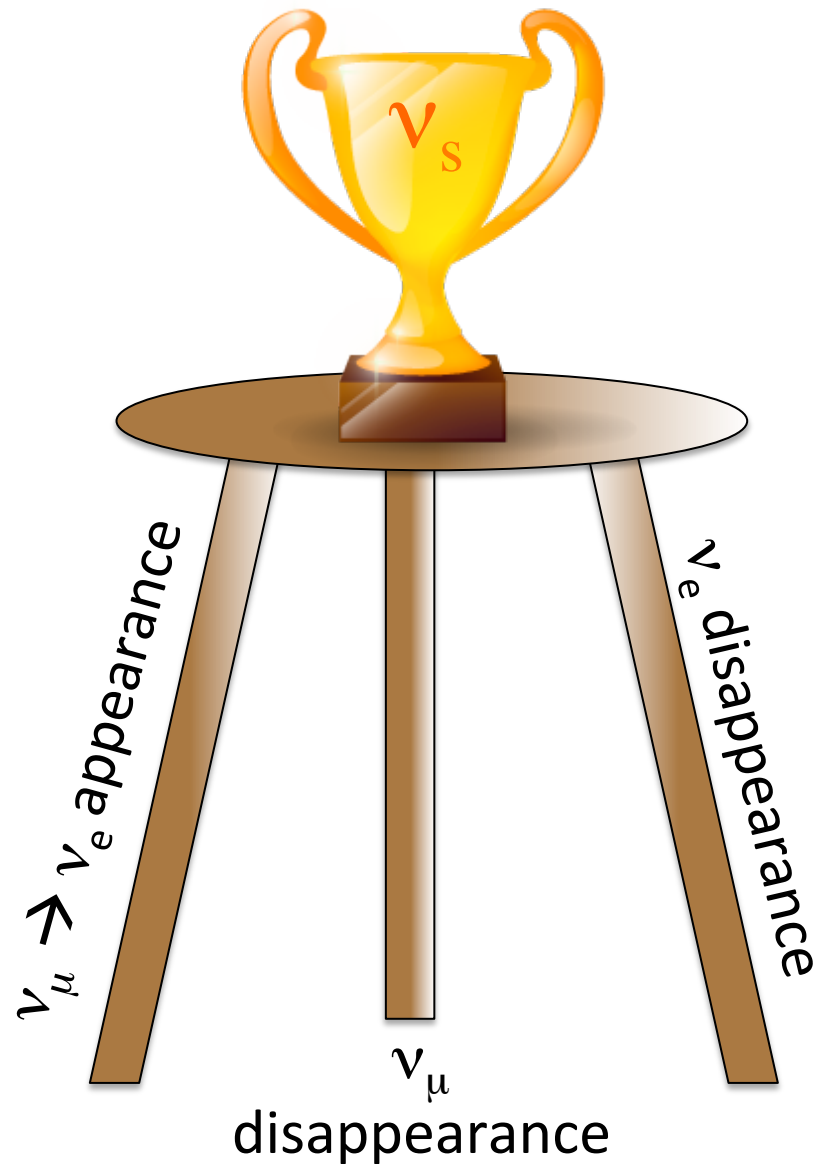
The long-term future of particle physics in the US is largely neutrino-based.

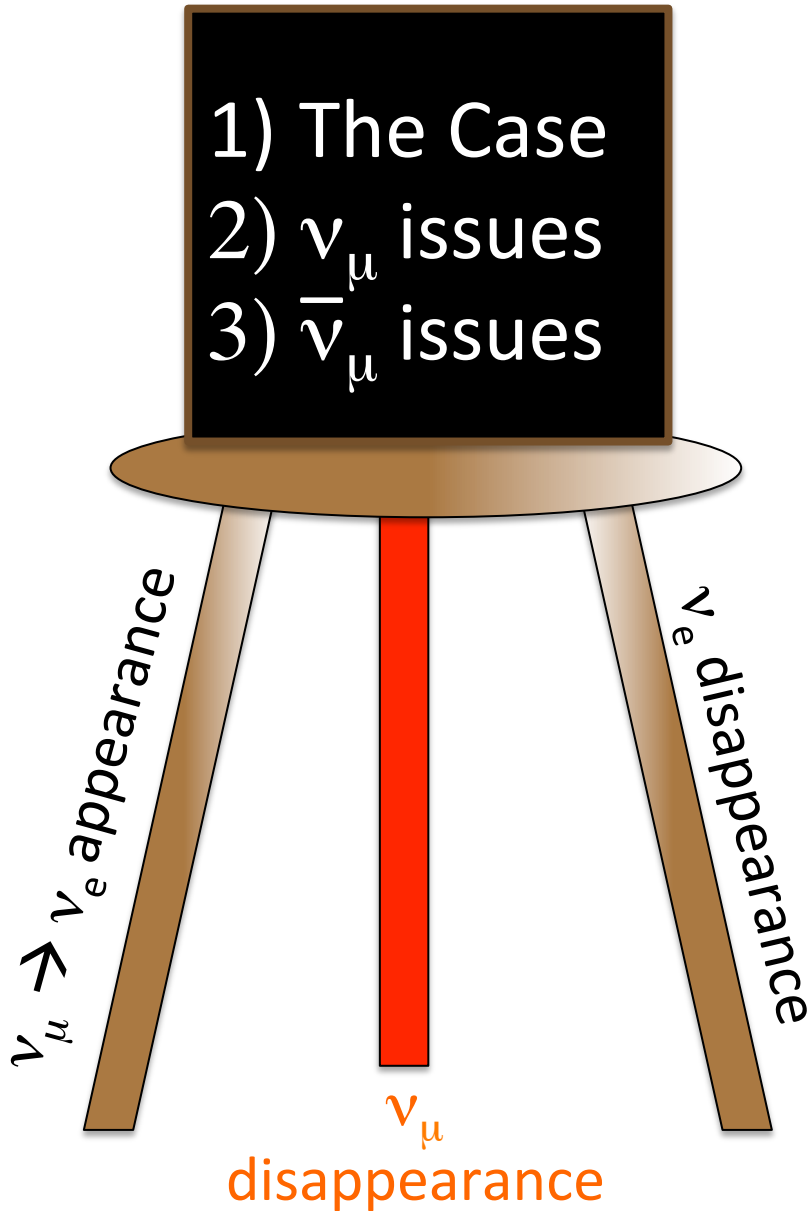
*...You get to the long-term future via the intermediate-term future!*

*An intermediate neutrino program that is scientifically decisive  
is the best path to long-term-future success  
for DOE and NSF.*



Success for sterile neutrinos studies  
sits on a “3-legged table”...



- 
- 1) The Case
  - 2)  $\nu_\mu$  issues
  - 3)  $\bar{\nu}_\mu$  issues

$\nu_\mu \rightarrow \nu_e$  appearance

$\nu_e$  disappearance

$\nu_\mu$

disappearance

# The Case for Near-Future $\mu$ -flavor Disappearance Experiments

There are “signals” in  $\nu_\mu \rightarrow \nu_e$  appearance and  $\nu_e$  disappearance,

No signal in  $\nu_\mu$  disappearance yet!  
...And there is “tension”

*Why?*

1. The observed signals are due to standard model processes, not sterile neutrinos
2. There are issues with some of the observed signals, but not others, leading to the wrong global fit.
3. The tension has to do with background, not signal, and  $\nu_\mu$  disappearance is right around the corner.
4. The tension has to do with more complicated physics than we are putting into our models.

## 1. What if all of the existing anomalies are Standard Model?

Given this as a scenario, should we not just wait on  $\nu_\mu$  disappearance  
Until we have more data from  $\nu_\mu \rightarrow \nu_e$  appearance and  $\nu_e$  disappearance?

I don't think so...

There are now many viable models that include steriles.

→ If we have not found them already, they still have a chance of being out there.  
*Probably as good of a chance as supersymmetry or axions.*  
We should be looking for them.

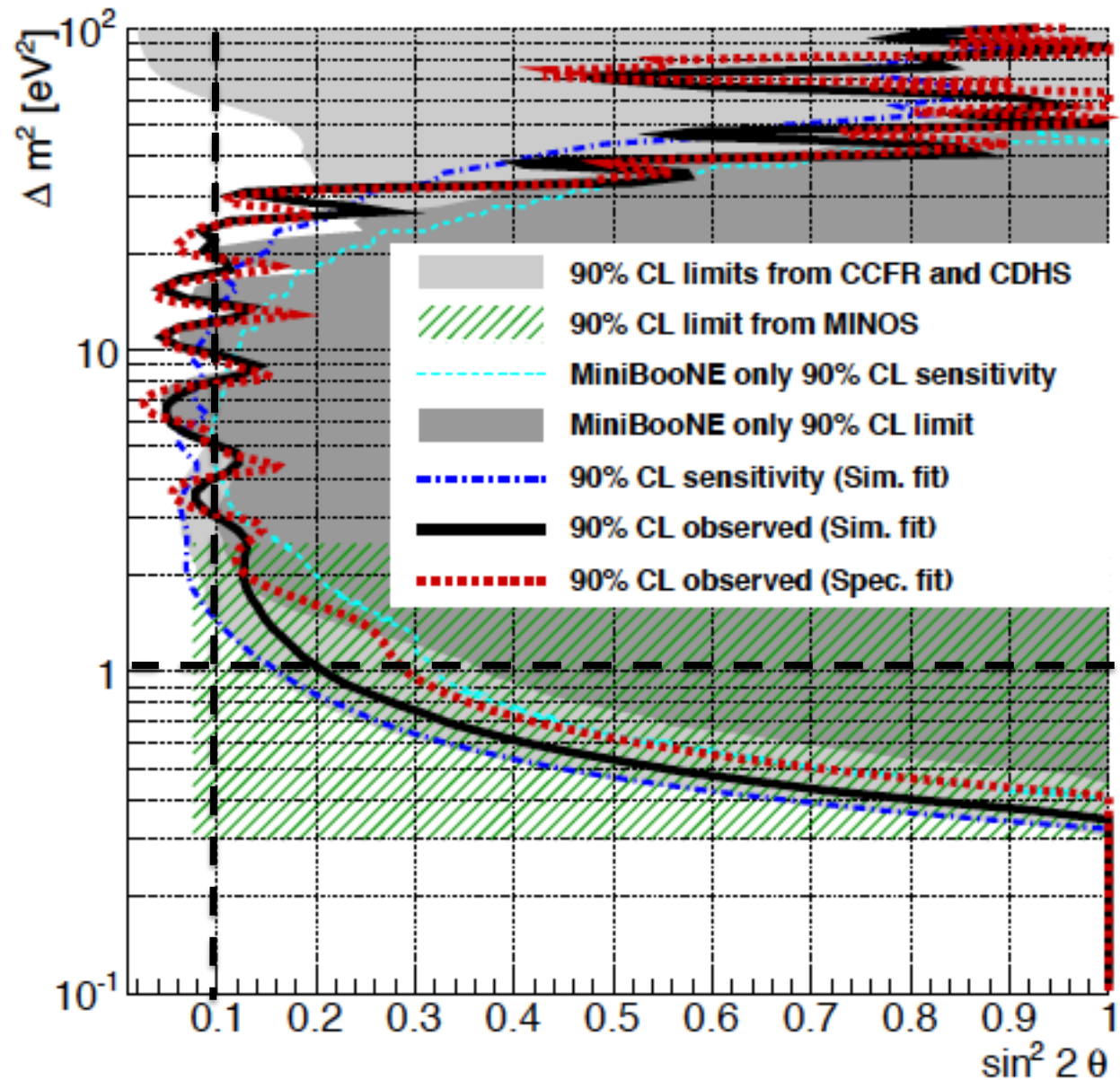
We should move forward with  $\nu_\mu$  disappearance experiments  
regardless of what happens with the present anomalies.



**In this event, we want the experiments to substantially increase  
the explored space past the present point where present data sits.**

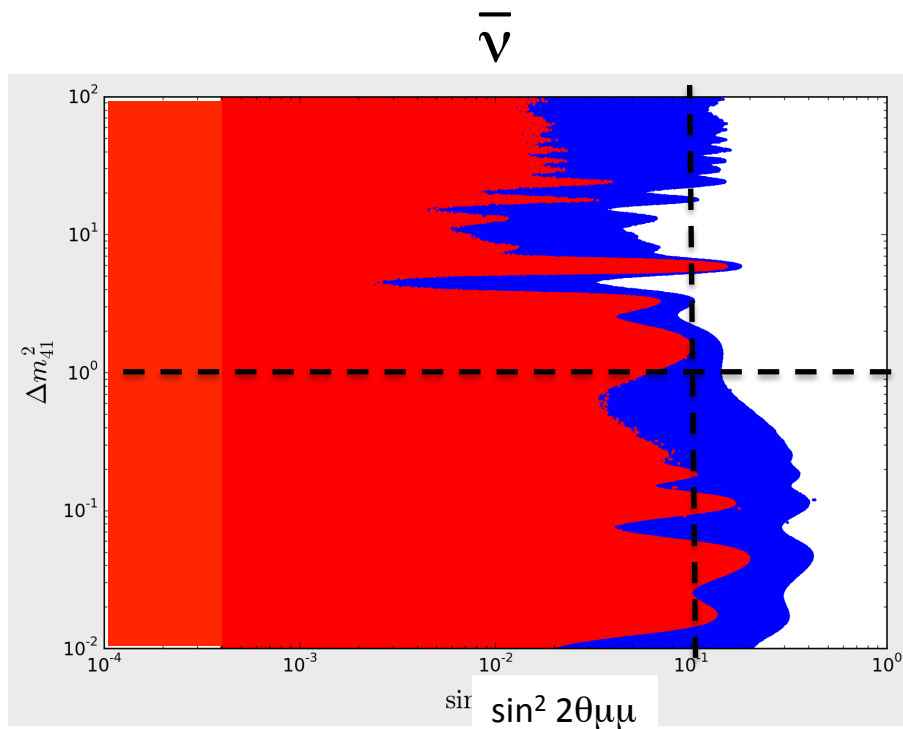
Existing experiments have rather similar limits!

We need  
substantial  
coverage  
Well below  
 $\sin^2 2\theta = 0.1$



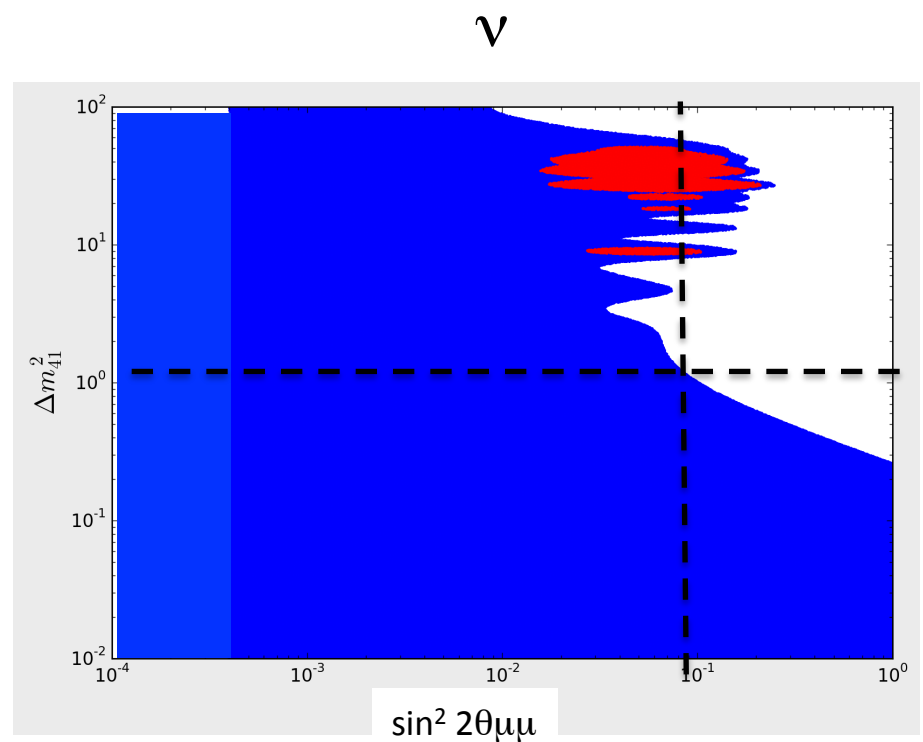


Putting all  $\mu$ -flavor disappearance data sets together,  
but separating the running modes...



99% CL, 90% CL

*Both modes need new experiments  
To give a substantial push below  $\sin^2 2\theta = 0.1$*



A mix of the experiments,  
especially CDHS+Sci/MiniBooNE,  
gives an allowed "island" but at  $< 2\sigma$

Fits by G. Collin, et al

2. What if we find there are issues with some of the observed signals,  
but not others, leading to the wrong global fit results?

Why suggest this?

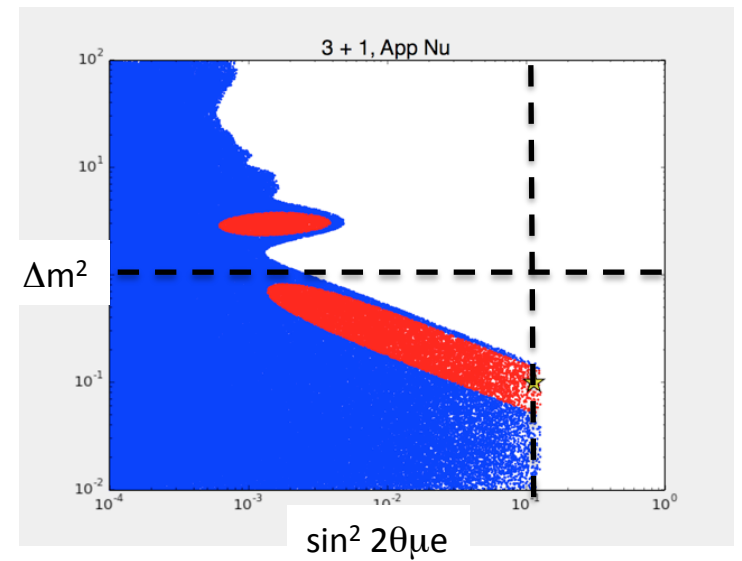
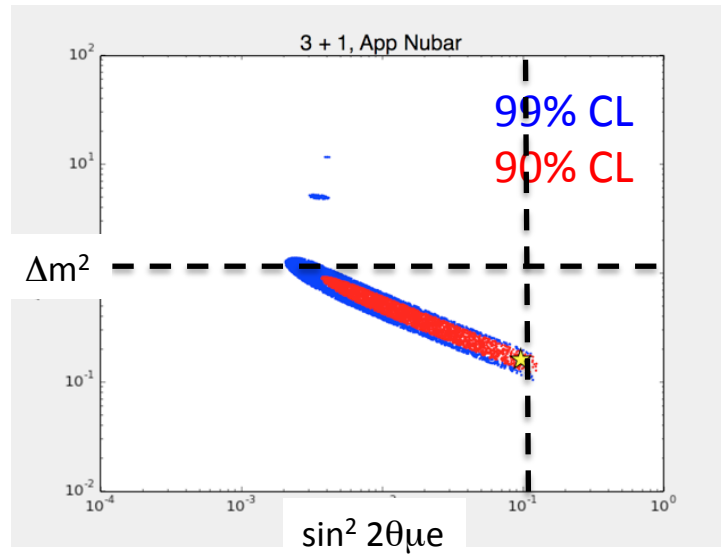
Let's divide the all of the data into two separate sets...  $\bar{\nu}$  and  $\nu$ ,

For each of these, let's do a 3+1 fit and look at the result  
(This is very close to a 3+2 fit and it is easy to visualize)

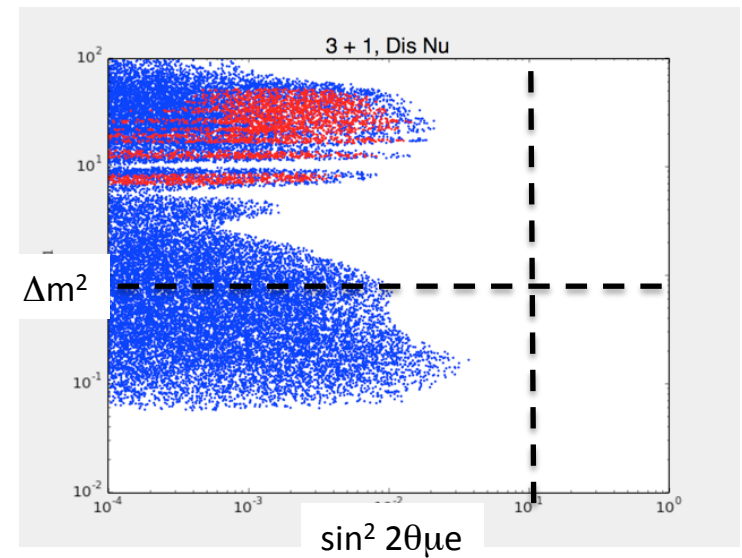
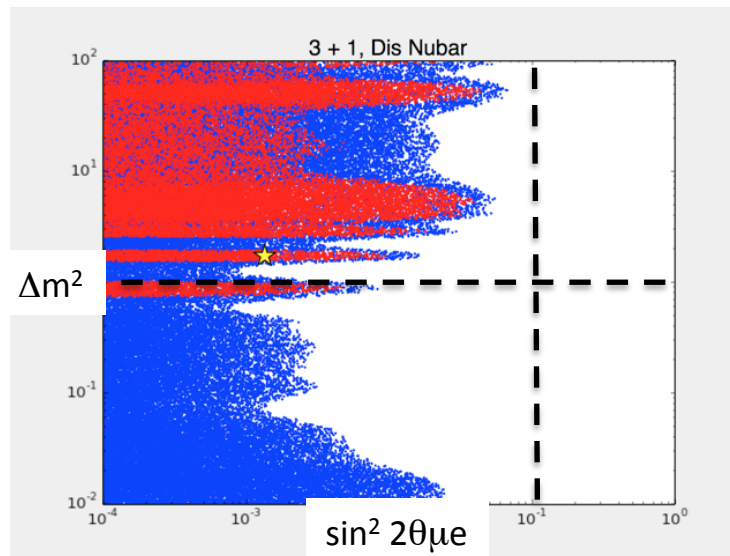
$\bar{\nu}$  3+1

$\nu$  3+1

Appearance



Disappearance

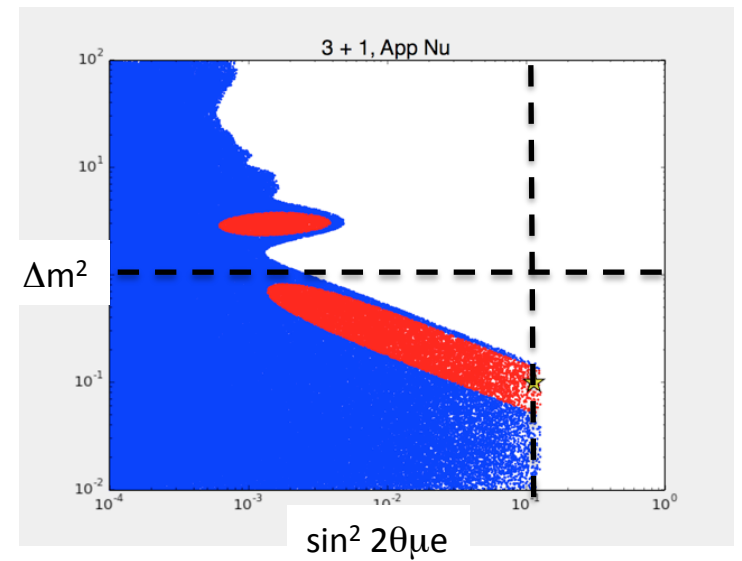
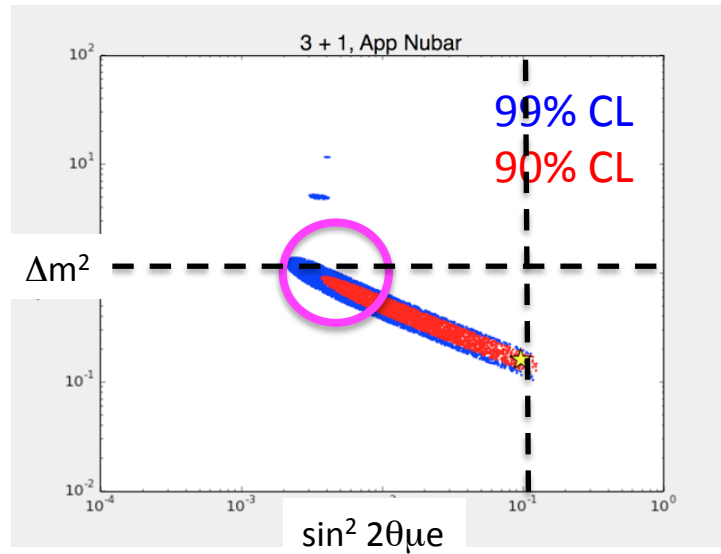


Fits by G. Collin, et al

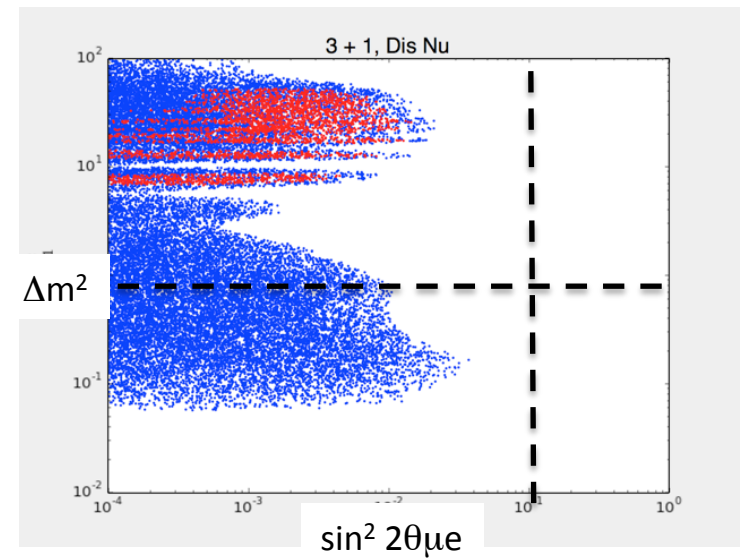
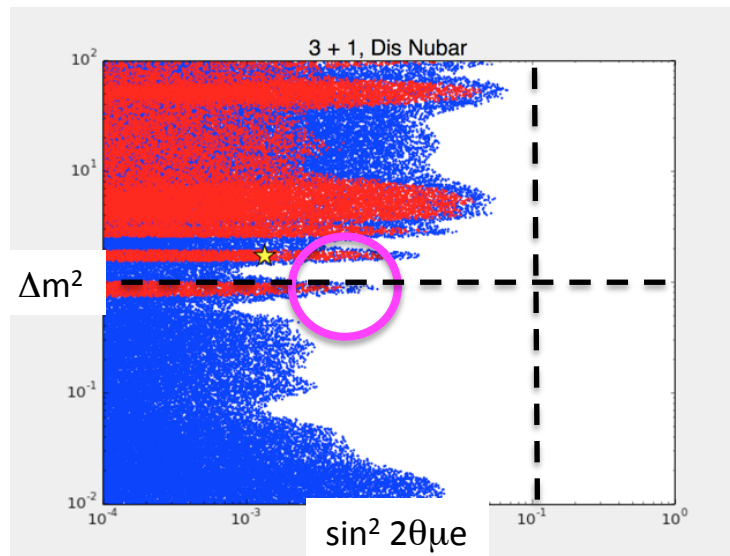
$\bar{\nu}$  3+1

$\nu$  3+1

Appearance



Disappearance

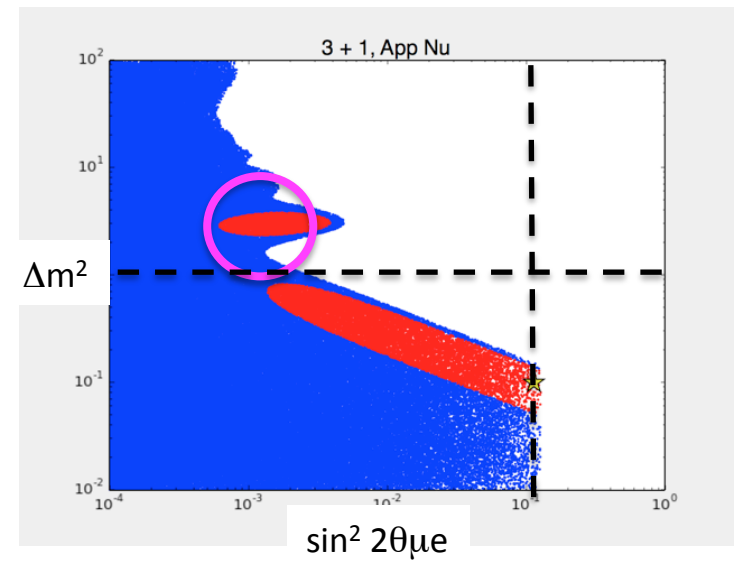
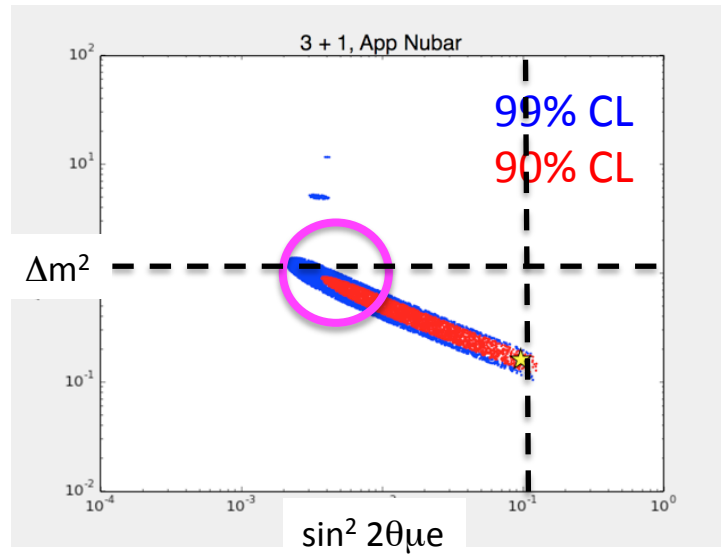


Fits by G. Collin, et al

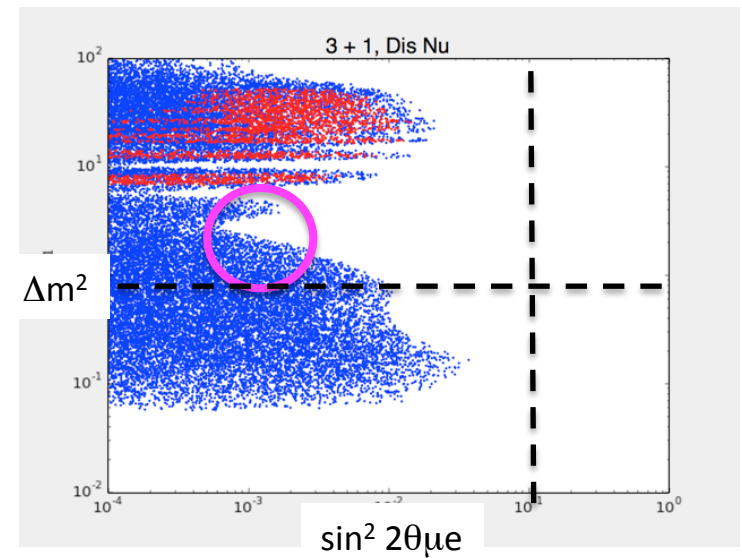
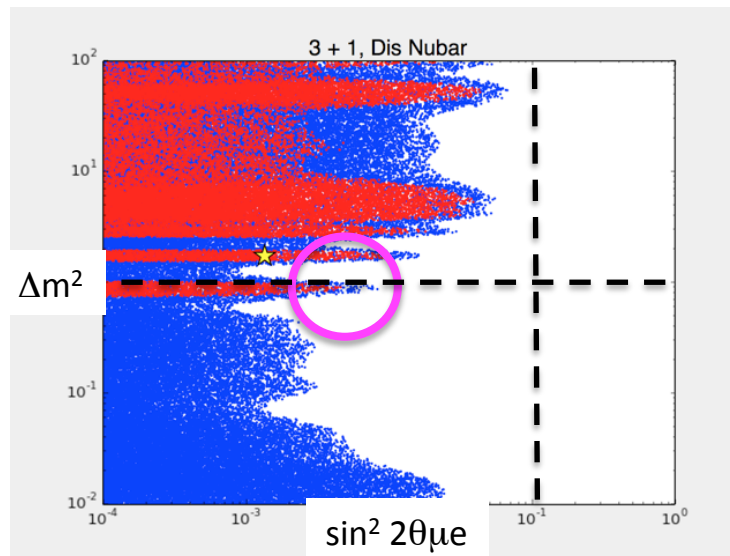
$\bar{\nu}$  3+1

$\nu$  3+1

Appearance



Disappearance



Fits by G. Collin, et al

All 3 types of searches in BOTH running modes need to be rechecked!

Even with only subsets being right,  
You are going to need a signal in  $\nu_\mu$  disappearance to make a sensible model!

It makes sense to proceed with new  $\nu_\mu$  disappearance searches now,  
even in this scenario.

3. What if the tension has to do with background, not signal, and  $\nu_\mu$  disappearance is right around the corner?

In this case it is obvious we should proceed with  $\nu_\mu$  disappearance experiments, but why would I even bring this question up?

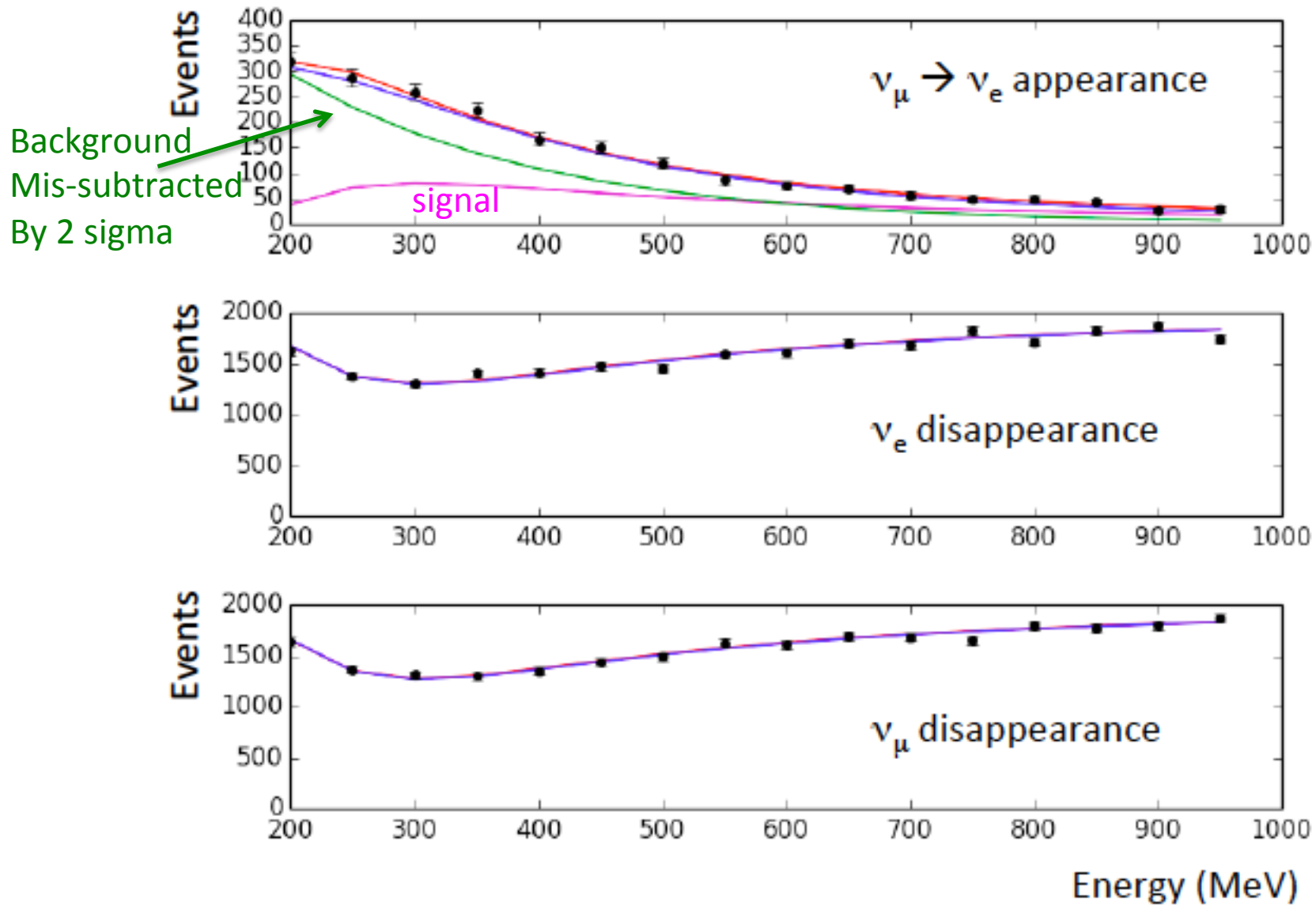
The PG test has a very specific flaw that can be demonstrated with a toy model.

In the case where...

- You have an underlying systematic effect that is not Gaussian, (like a background that is mis-subtracted)
- And the associated shape correlates with a signal somewhere in parameter space,
- The test will return tension, even though the global fit is correct.

→ This is true even when the mis-subtraction is small when measured in “sigmas”

Consider this model





What the PG Test evaluates:

$$\chi_{PG}^2 = \chi_{glob}^2 - (\chi_{app}^2 + \chi_{dis}^2)$$



Because of the correlation between the shape of the mis-subtracted background and the signal, this fit goes very wrong.

The result is that random experiments find substantial “tension” but in fact the global fit is returning exactly the toy model parameters.

We are in a situation where the global  $\chi^2$  is good. but the PG returns tension.

We may be in the same situation as this toy model...

$\nu_\mu$  disappearance may be right around the corner!

	$\chi_{best}^2$	Pr
Null	290	1.5%
3+1	240	43%
3+2	235	45%

#### 4. What if Nature is more complex than we thought?

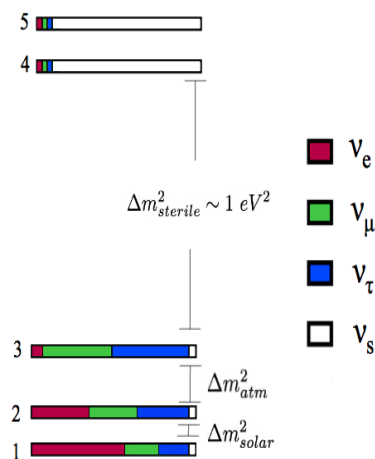
A 3+1 model is clearly ridiculously naïve.  
 We should NOT be using 3+1 models as tests.  
 They are very misleading as to the capability of an experiment.

But just how complex is nature?

Is it a 3+2 model? This has a lot of obviously strange assumptions too...

The only parameters are...

$$\Delta m_{41}^2, \Delta m_{51}^2, |U_{e4}|, |U_{\mu 4}|, |U_{e5}|, |U_{\mu 5}|, \Phi_{45}$$



We are missing the other  $\Delta m^2$ s and  
 The light neutrino CP violating term!

And 3+3, as complex as those fits are,  
 we are also missing a lot of terms!

But maybe Nature is even more complex!

What if sterile neutrinos can also decay?

100% decay is ruled out as a model, But 3+N + decay is definitely not ruled out!

What if there are non-standard interactions affecting all of our measurements?

What if there is CPT Violation?

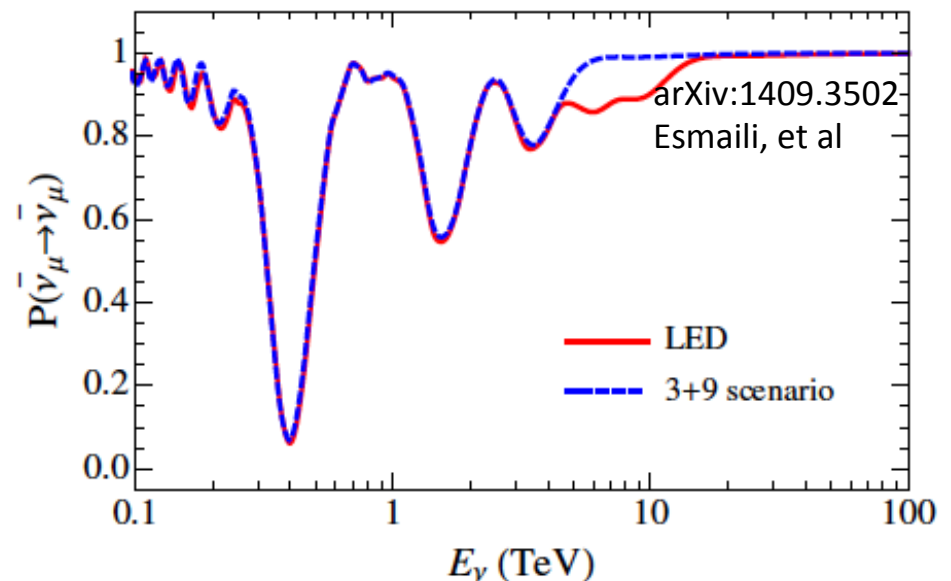
What if there are large extra dimensions? →

What if...

What if...

What if..

Your favorite theory  
extension here!



We really need  $\nu_\mu$  disappearance experiments where we reconstruct the oscillation wave!

It also says we need better modeling tools...



# SQuIDS/nu-SQuIDS

*C.A. Argüelles Delgado et al. [arXiv:1412.3832]*

## What is it?

Is a software framework written in C++ that **evolves quantum mechanical ensembles**.  
nu-SQuIDS **calculates neutrino propagation** (oscillation+interactions).

## What can it do?

- ❑ Calculate neutrino oscillation probabilities in 3 generations (can configure mixing angles, CP phases, and mass splittings).
- ❑ Ready to use in: short baseline, long baseline, atmospheric, and solar neutrino oscillation experiments.
- ❑ Incorporates neutrinos' non-coherent interactions (includes tau regeneration).
- ❑ Can handle collective neutrino interactions (e.g. super nova), as well as neutrino-antineutrino interactions.
- ❑ Easily extendable to BSM physics scenarios. **Sterile neutrinos**, NSI, and LV already implemented!

## Get it here:

<https://github.com/jsalvado/SQuIDS>

<https://github.com/arguelles/nuSQuIDS>

Please  
Support Neutrino  
Phenomenology!



1. What if all of the existing anomalies are Standard Model?

2. What if we find there are issues with some of the observed signals, but not others, leading to the wrong global fit results?

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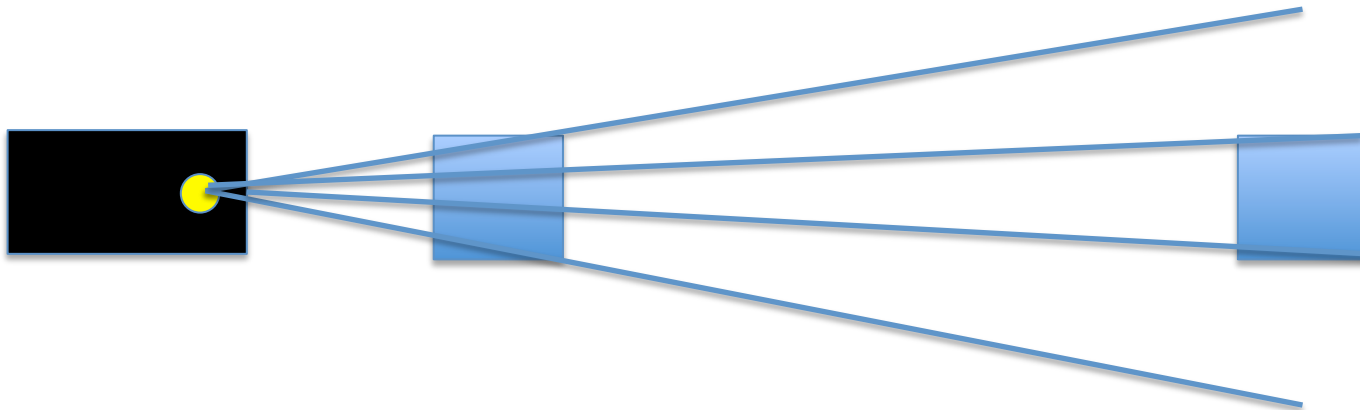
## The Bottom Line for This Case:

All 4 of these scenarios point to the need for  
*much higher precision  $\nu_\mu$  disappearance experiments*  
in neutrino and antineutrino mode  
in the near future

$\nu_\mu$  Disappearance  
Experiments  
Need Well-understood  
 $\nu_\mu$  Fluxes

Decay-in-flight fluxes are difficult to simulate very well.

So most  $\nu_\mu$  disappearance experiments use near-far ratios.

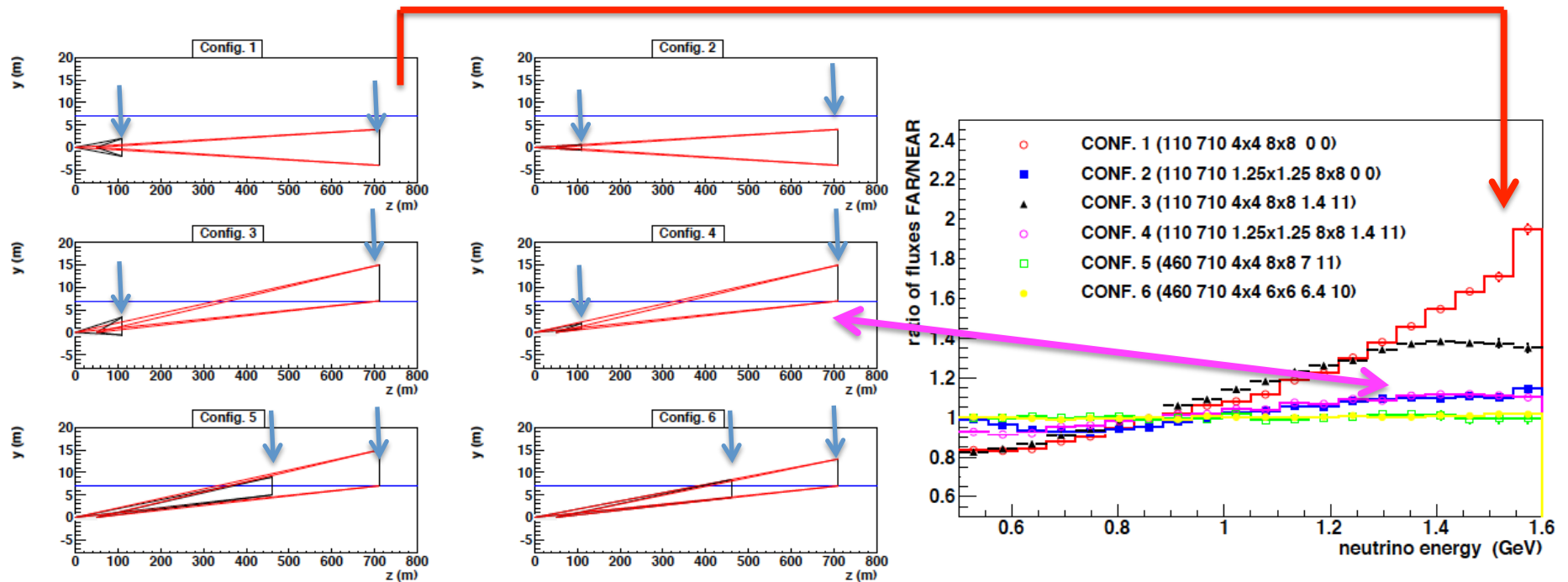


With CC interactions, you can reconstruct the neutrino direction,  
And select only those events in the near detector that project to the far detector.

# A study of configurations for the BNB (FNAL) beam by the NESSiE Collaboration

hep-ex/1404.2521

Flux ratios can show big differences  
Between near and far detectors!



4 m x 4 m near  
(Both designs have an 8 m x 8m far detector)

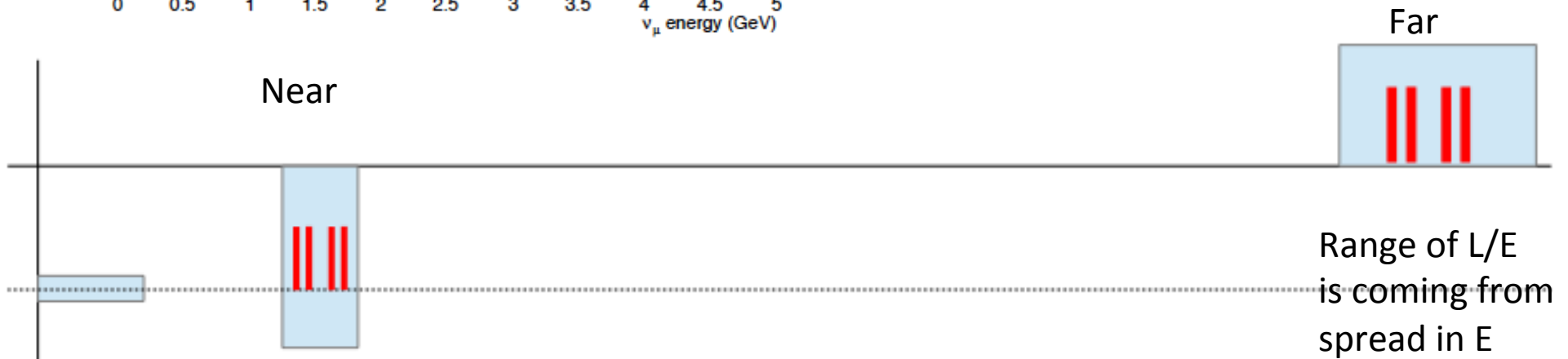
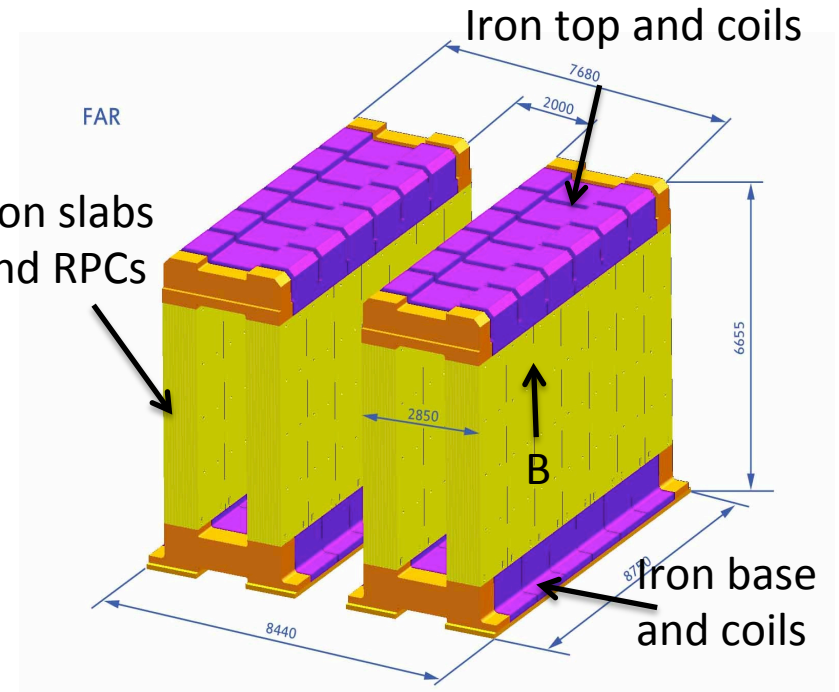
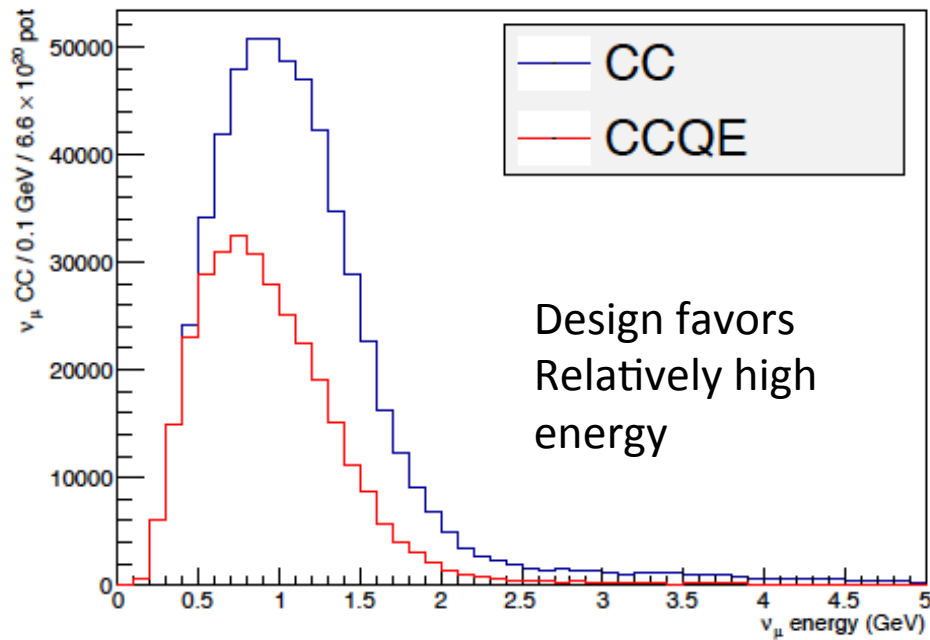
1.25 m x 1.25 m near

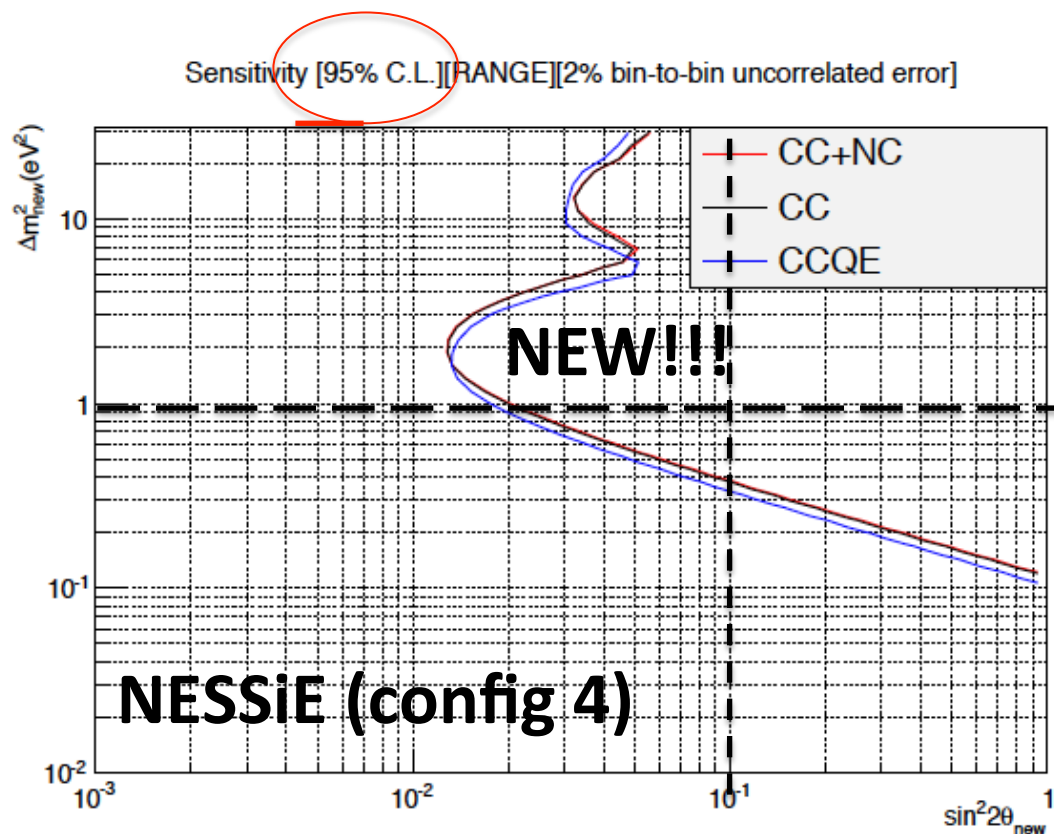
Choice is config. 4

Need to understand the  
shape to ~1% level!

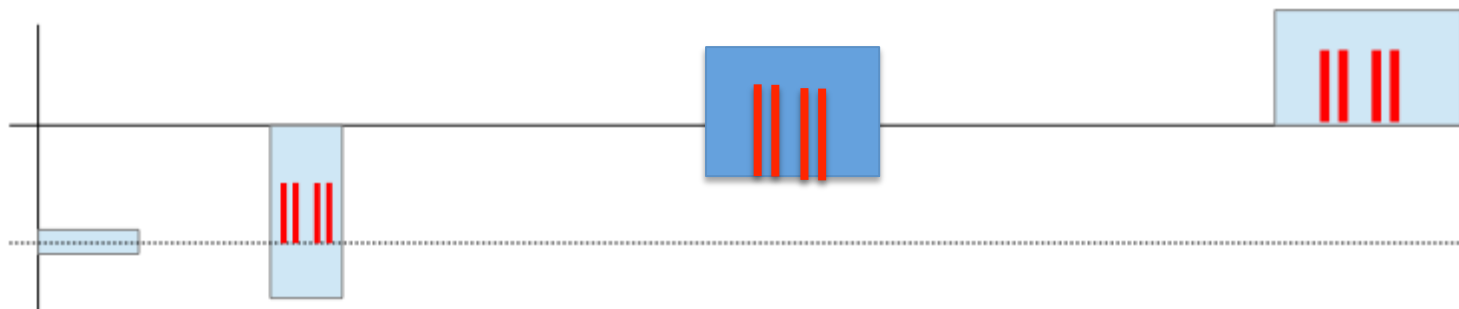


NESSiE is a magnetized detector,  
With spectrometer design (and components!) drawn from OPERA





Really great sensitivity, but it will only be a single point measurement in L. What if this experiment had an intermediate detector too?



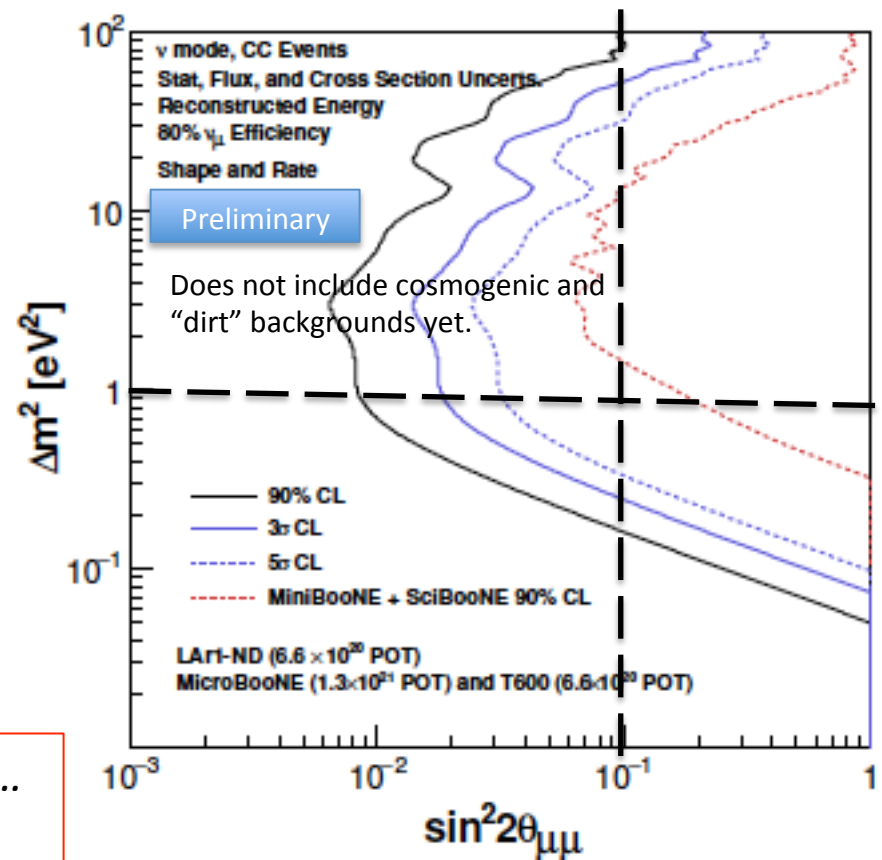
NESSiE is new to the plan....

Why not just rely on the “SBN” liquid argon detectors and forget other approaches?  
After all, they will already be there for the  $\nu_e$  appearance studies!

LAr technology is not  
optimized for muon reconstruction.

Exiting muon momentum is determined  
through multiple scattering.

Very soon MicroBooNE will give us a  
good sense of how well this can be done  
on the BNB line!



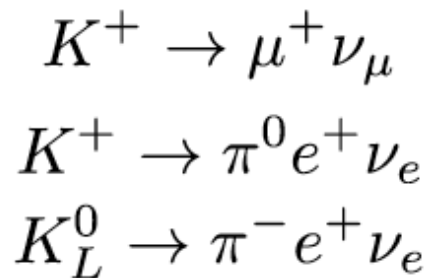
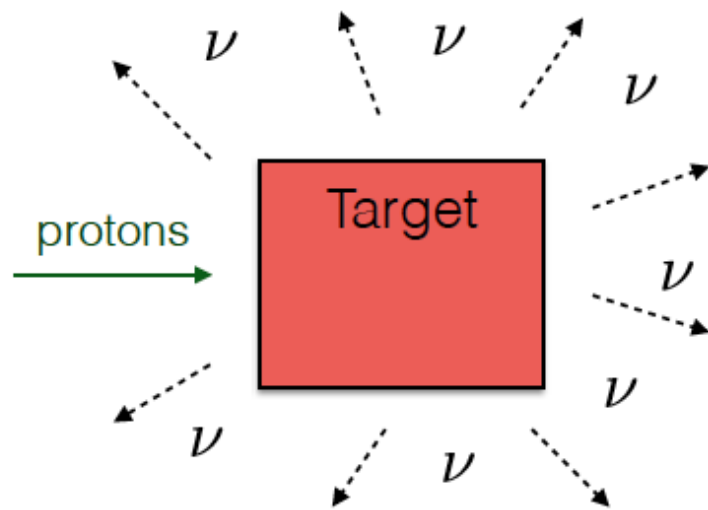
*It is great we have SBN-LAr, but...*

*If we really want to understand  
 $\nu_\mu$  disappearance waves,  
then we probably need detectors  
that specialize in muons*

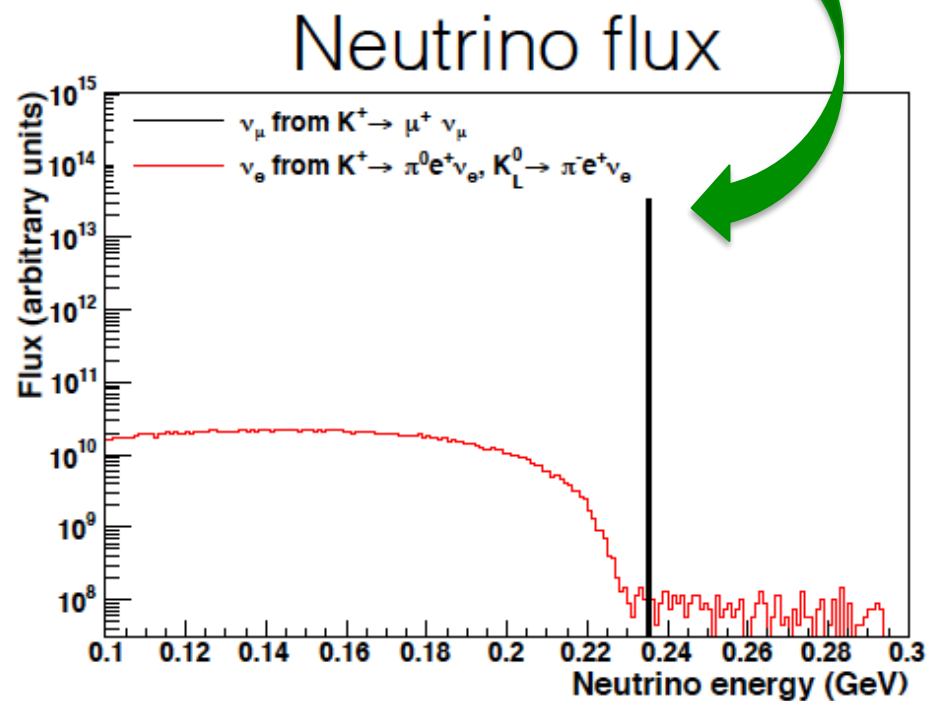


An idea that no one has looked at yet for  $\nu_\mu$  disappearance:  
Can we use KDAR?

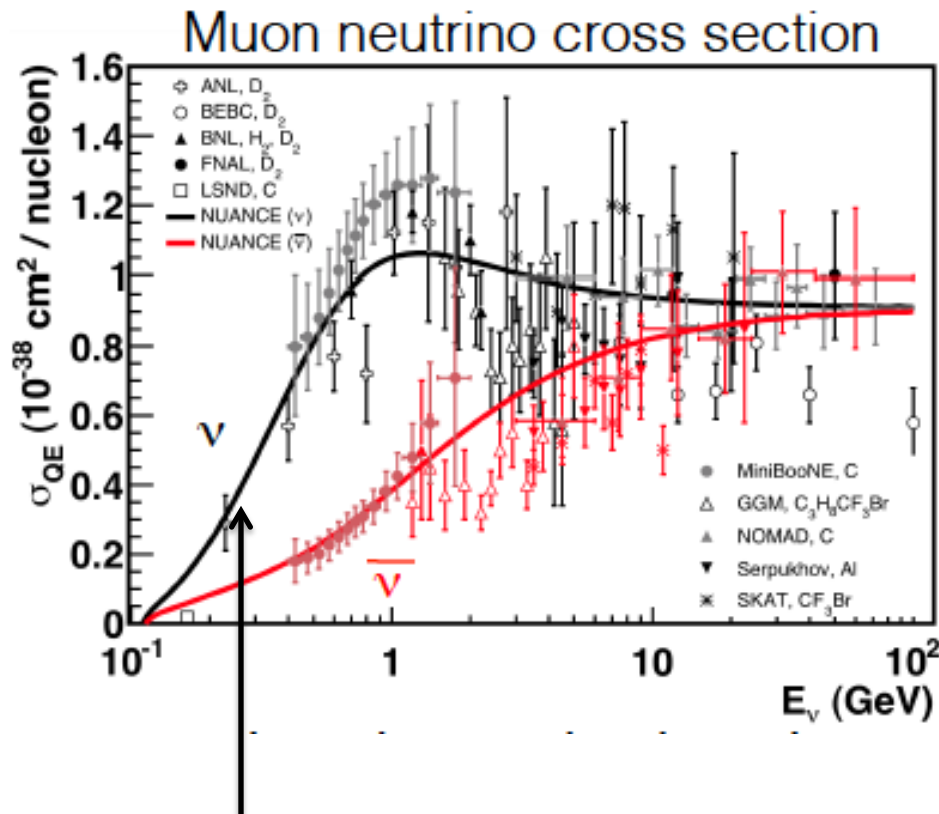
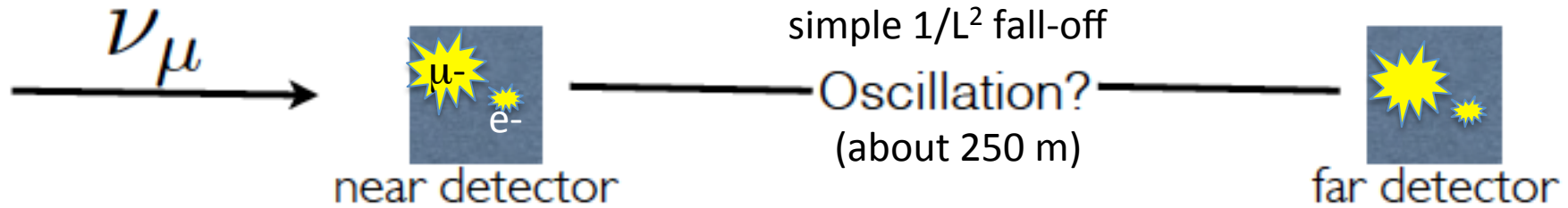
*J. Spitz, Phys. Rev. D 85 093020 (2012)*



Unlike the BNB line, this flux is mono-energetic and isotropic.



JPARC MLF is offering this beam NOW



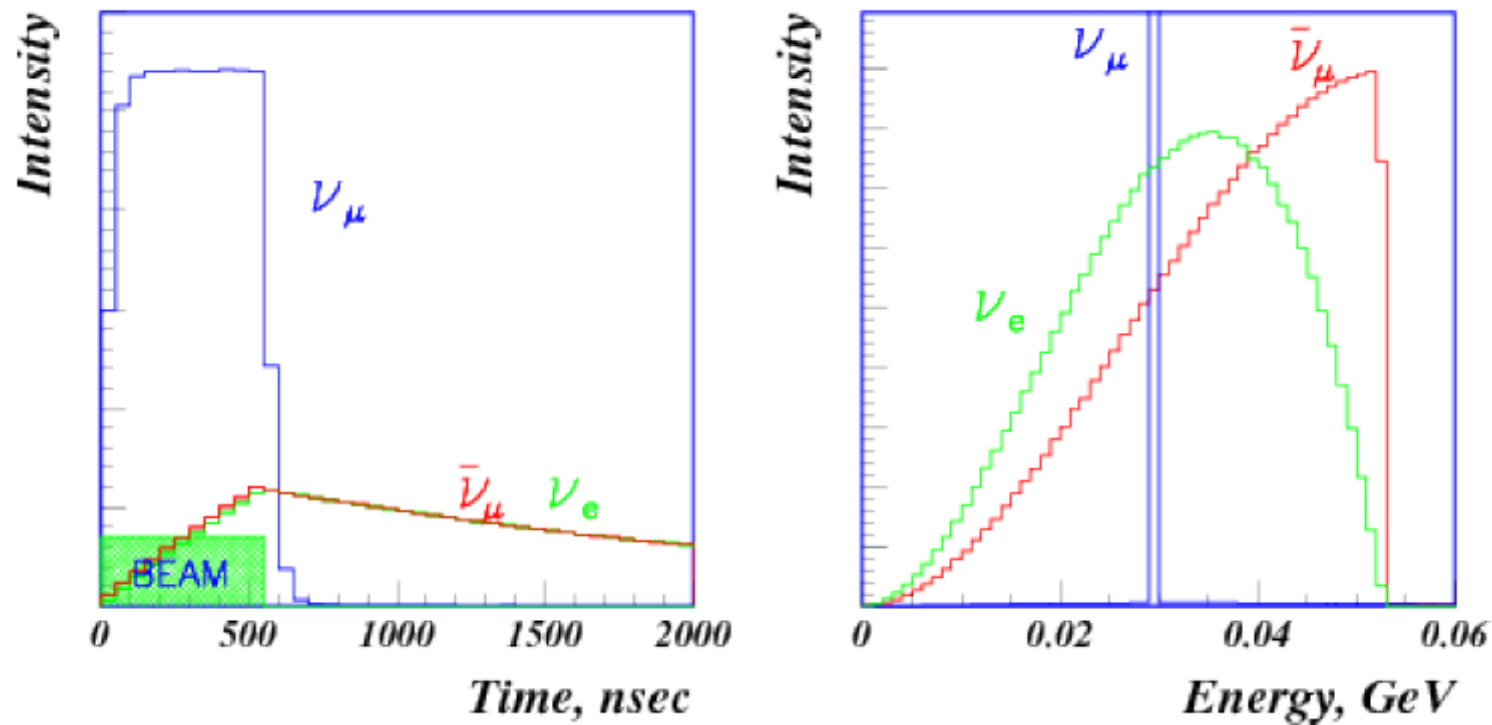
No sensitivity calculation yet,  
but seems to be  
a practical design  
**for tracing  
the osc wave.**

Just put lots of little 50 ton  
liquid scintillator  
detectors in line...  
<\$10M, fast, & Beam exists!

Despite low xsec @ JPARC MLF -  
1000  $\nu_\mu$  CC events at 250 m with 50 ton detector in 4 years

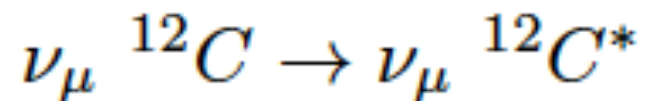


$\pi/\mu$  DAR also has a monoenergetic  $\nu_\mu$  flux,  
And, given a bunched beam structure, it is prompt!



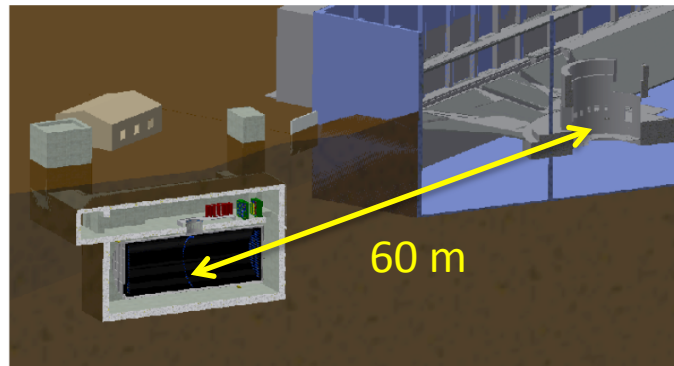
Too low energy for CC scattering,

OscSNS Concept, use NC:

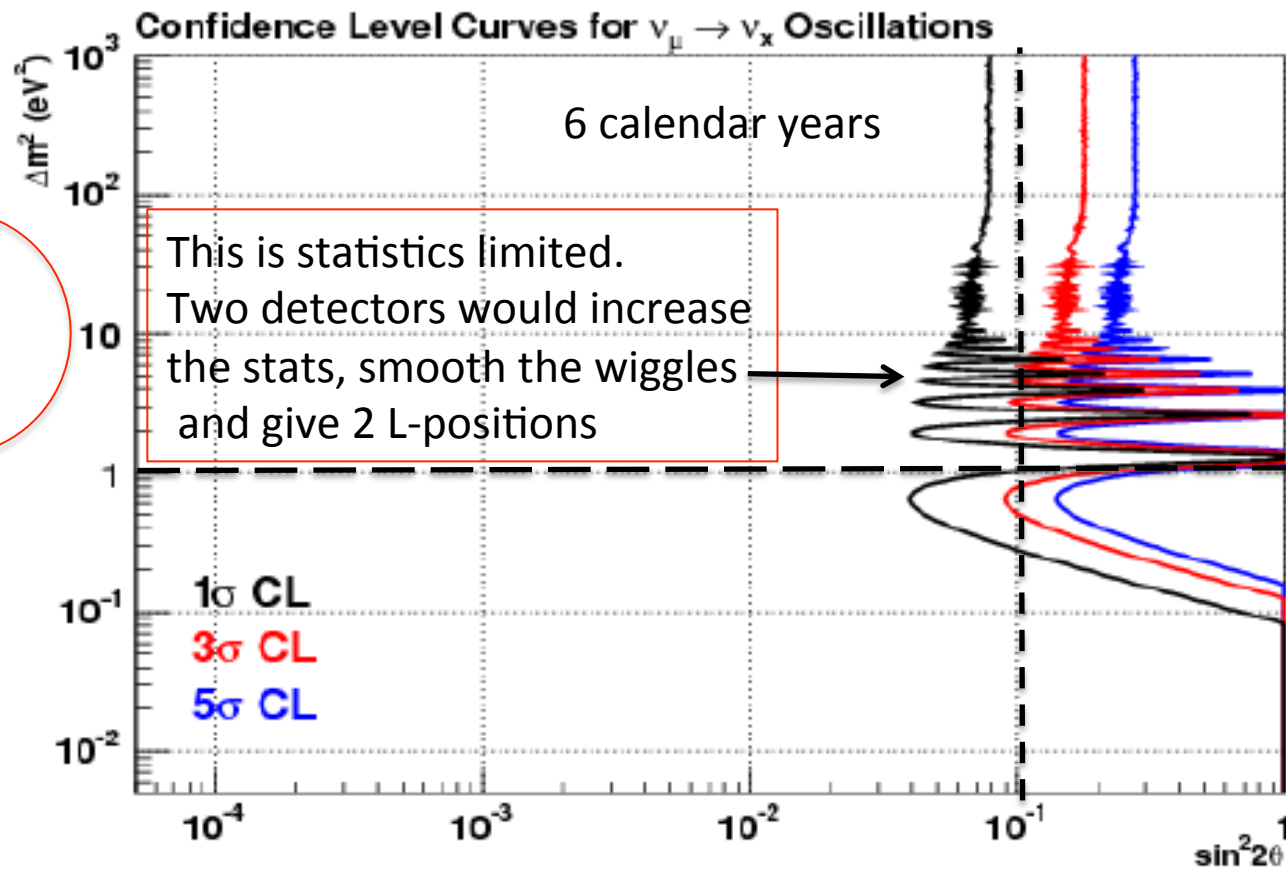


hep-ex/1305.4189

OscSNS is 800 tons  
Of lightly doped LS



1.4 MW  
source



$\overline{\nu}_\mu$  Disappearance Studies  
Suffer from a Lack of  
Pure Beams



Where can we find high rate, high purity muon antineutrino fluxes?

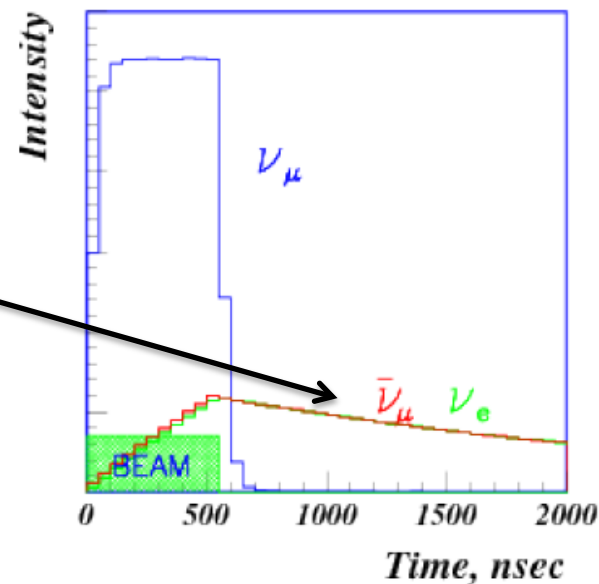
$\pi/K$  decay in flight (eg BNB)  $\rightarrow$  Very low rates. Poor purity. No present plans!

What about DAR? Can OscSNS do it?

$\rightarrow$   $\mu$ DAR mixes

$\nu_e$  and  $\bar{\nu}_\mu$  disappearance signals  
in NC mode

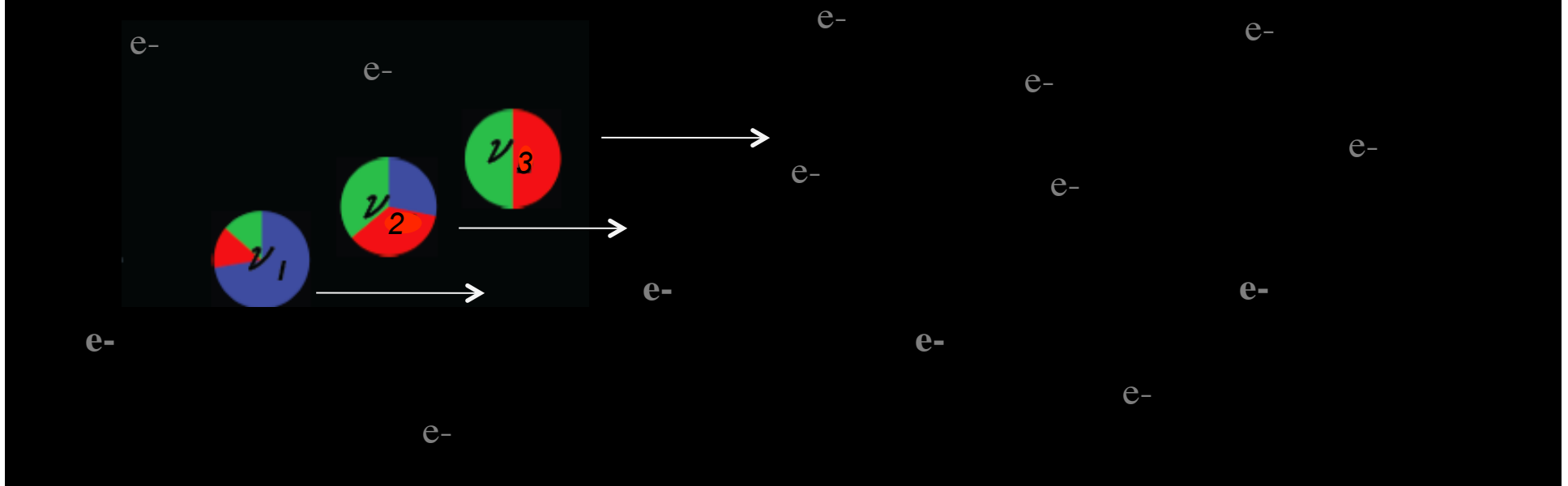
Atmospheric? Also a  $\nu/\bar{\nu}$  mix!



It looks really hard  
to find  $\bar{\nu}_\mu$  flux purity...

Thinking a little outside the box on this problem....

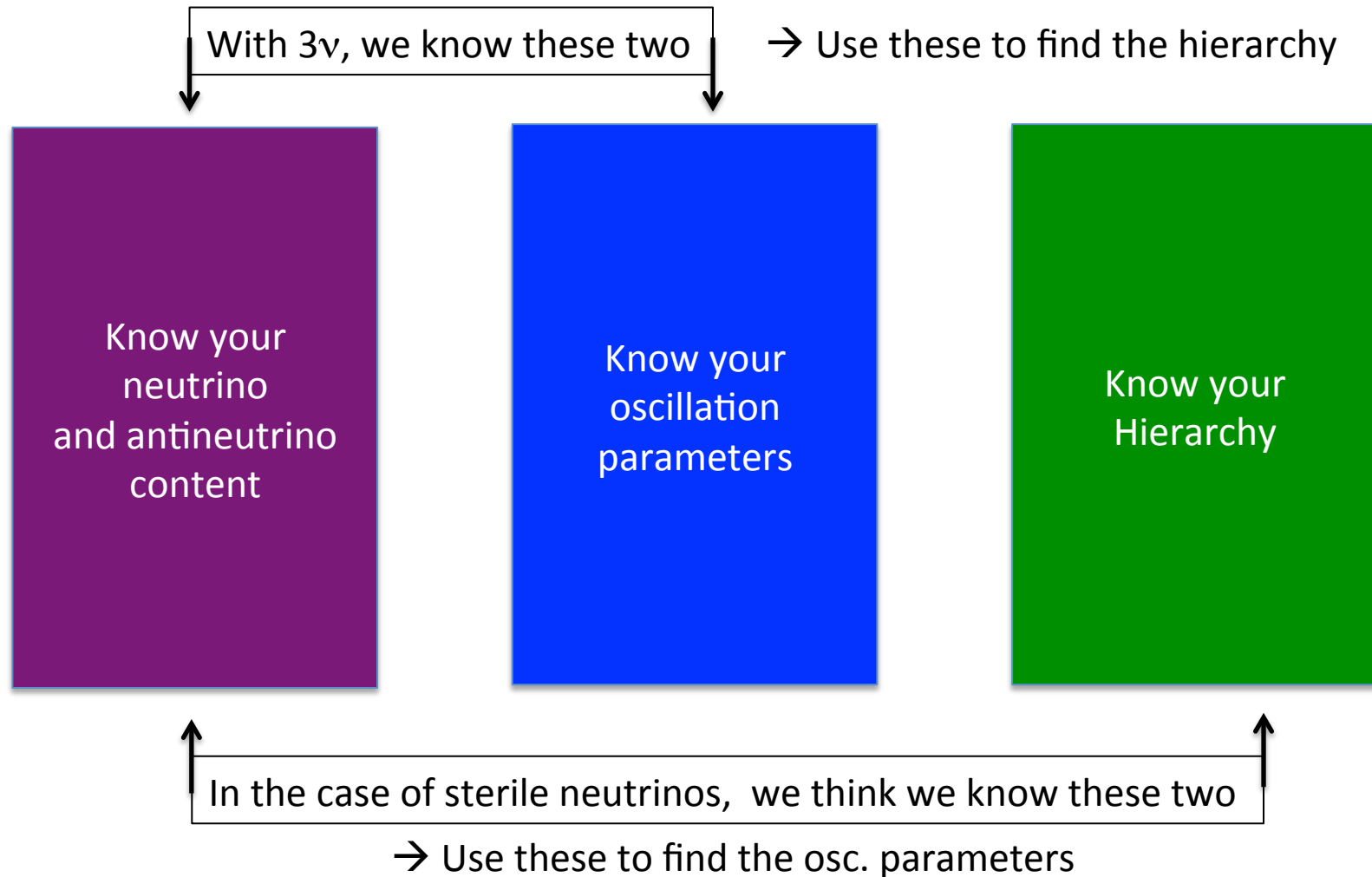
## “Matter Effects”



- Electrons in the earth produce a “weak-interaction potential”
- Effect is opposite sign for neutrinos and antineutrinos

At high energies and long distances, oscillations will be affected.  
→ You get BIG disappearance effects in only one mode ( $\nu$  or  $\bar{\nu}$ ),  
which mode depends on the hierarchy

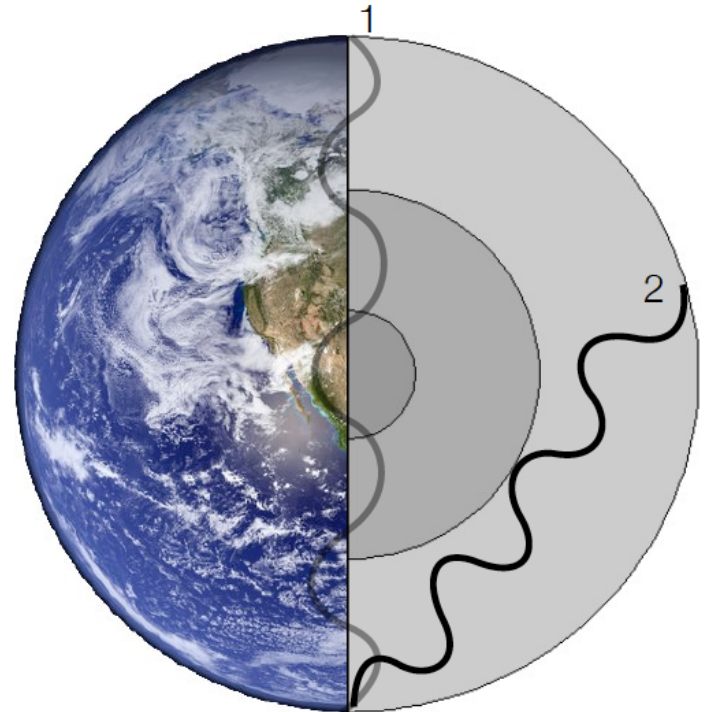
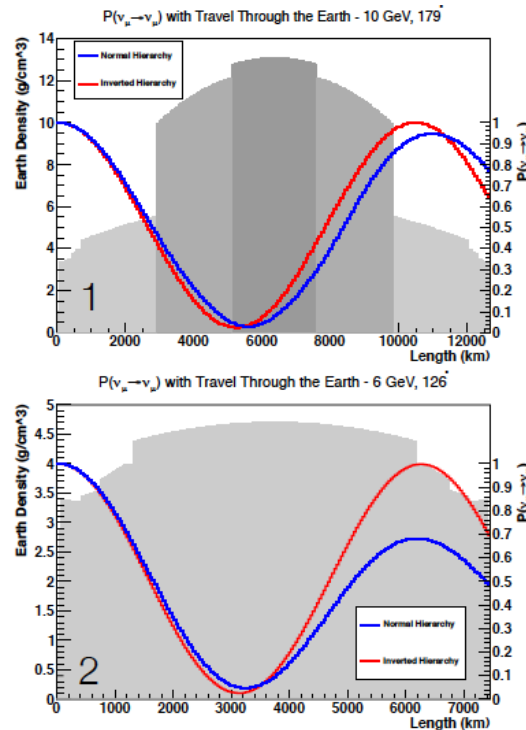
# Pick 2 -- Get the 3rd



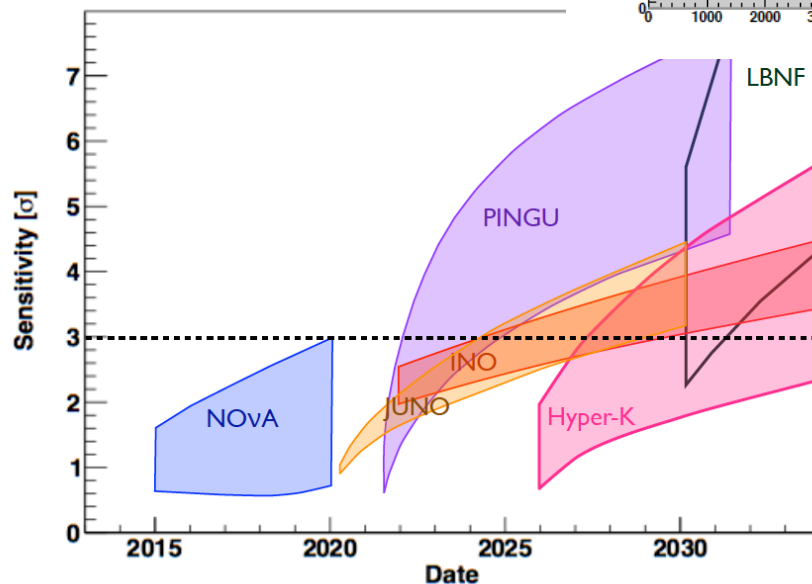
Let's use IceCube/PINGU as our example...

# This is how PINGU works for 3-Neutrino Hierarchy Studies...

Because the oscillation parameters are known, the waves through the paths are predictable



Graphic from J. Koskinen

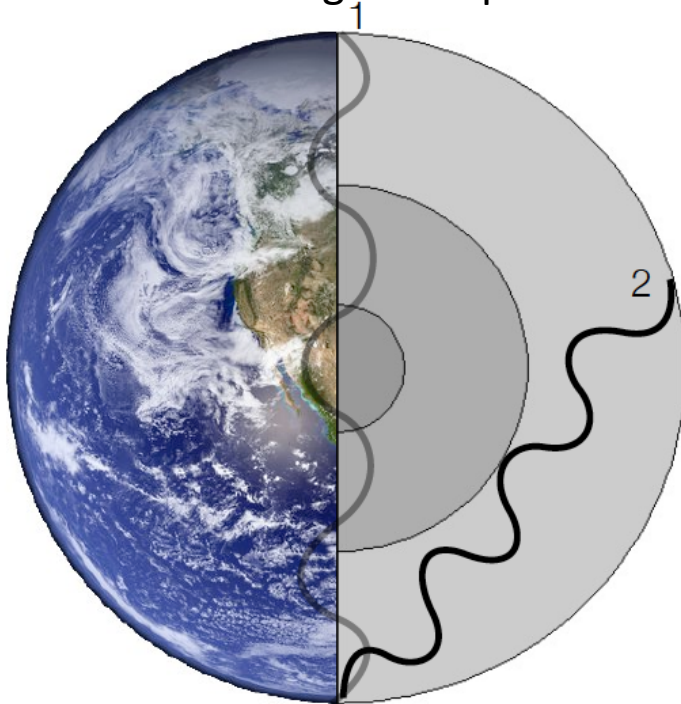


The result is very powerful sensitivity to the hierarchy!

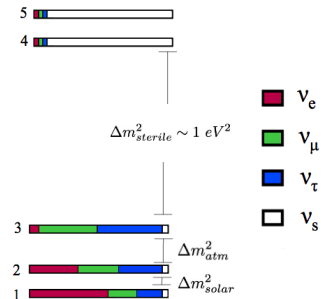


Using the exact same trick for sterile neutrinos, knowing the hierarchy, not the osc parameters...

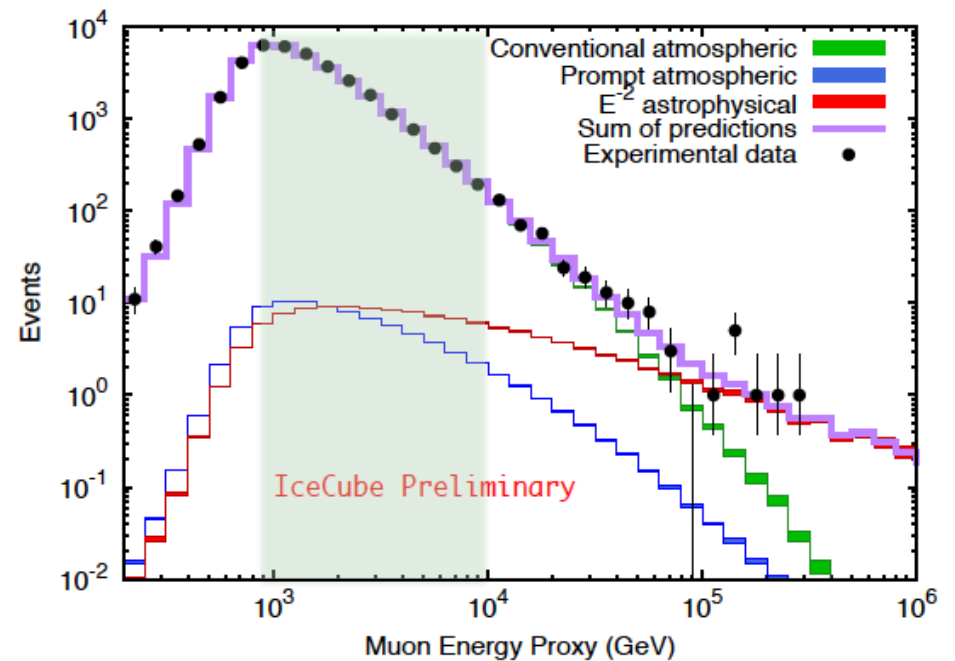
Use the angular dependence



Use the Hierarchy:



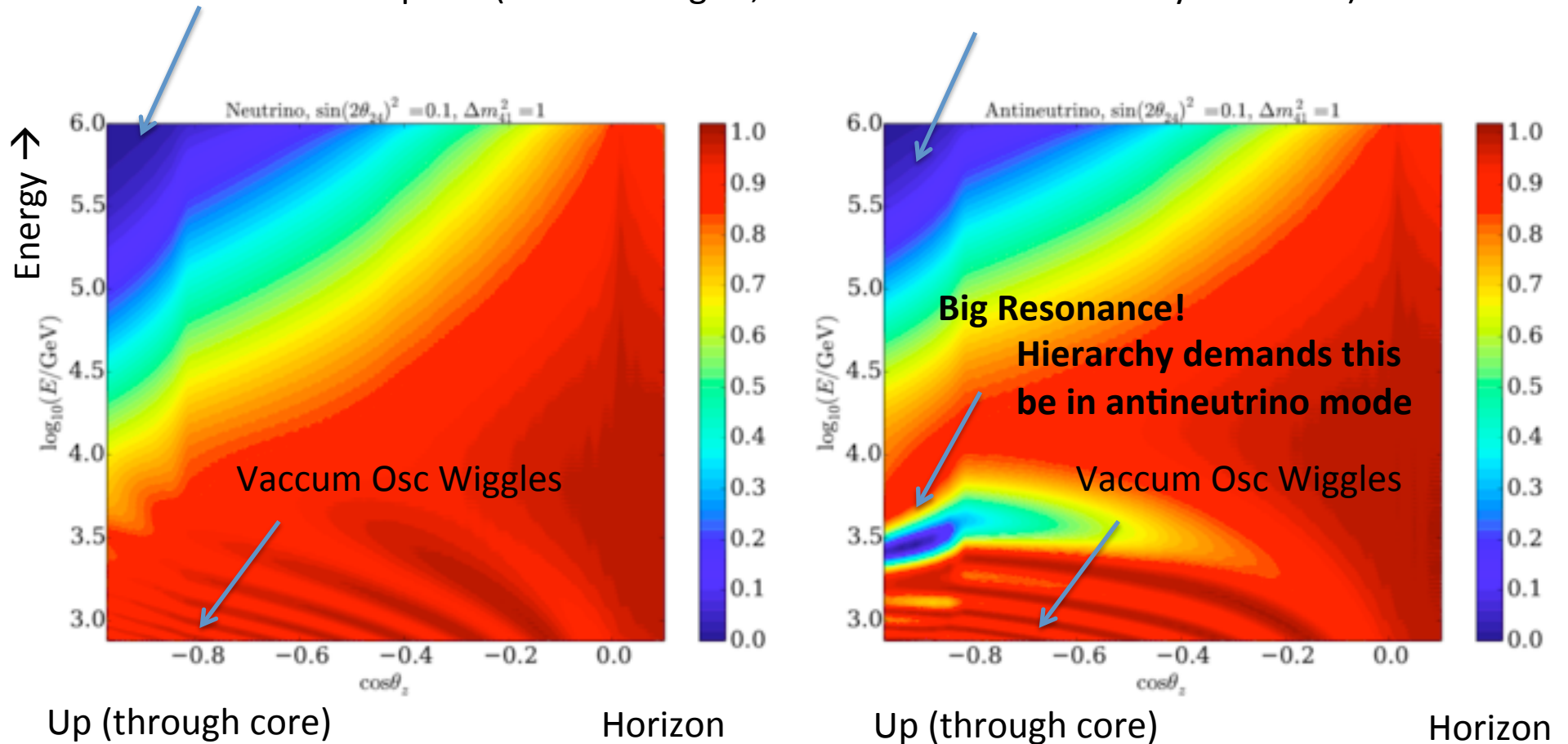
Neutrinos in the 1 to 10 TeV range (optimal for sterile neutrino matter effects In the earth)



This is where IceCube has enormous statistics!

# Oscillograms 101

Neutrino Absorption (at TeV energies, neutrinos are absorbed by the earth)

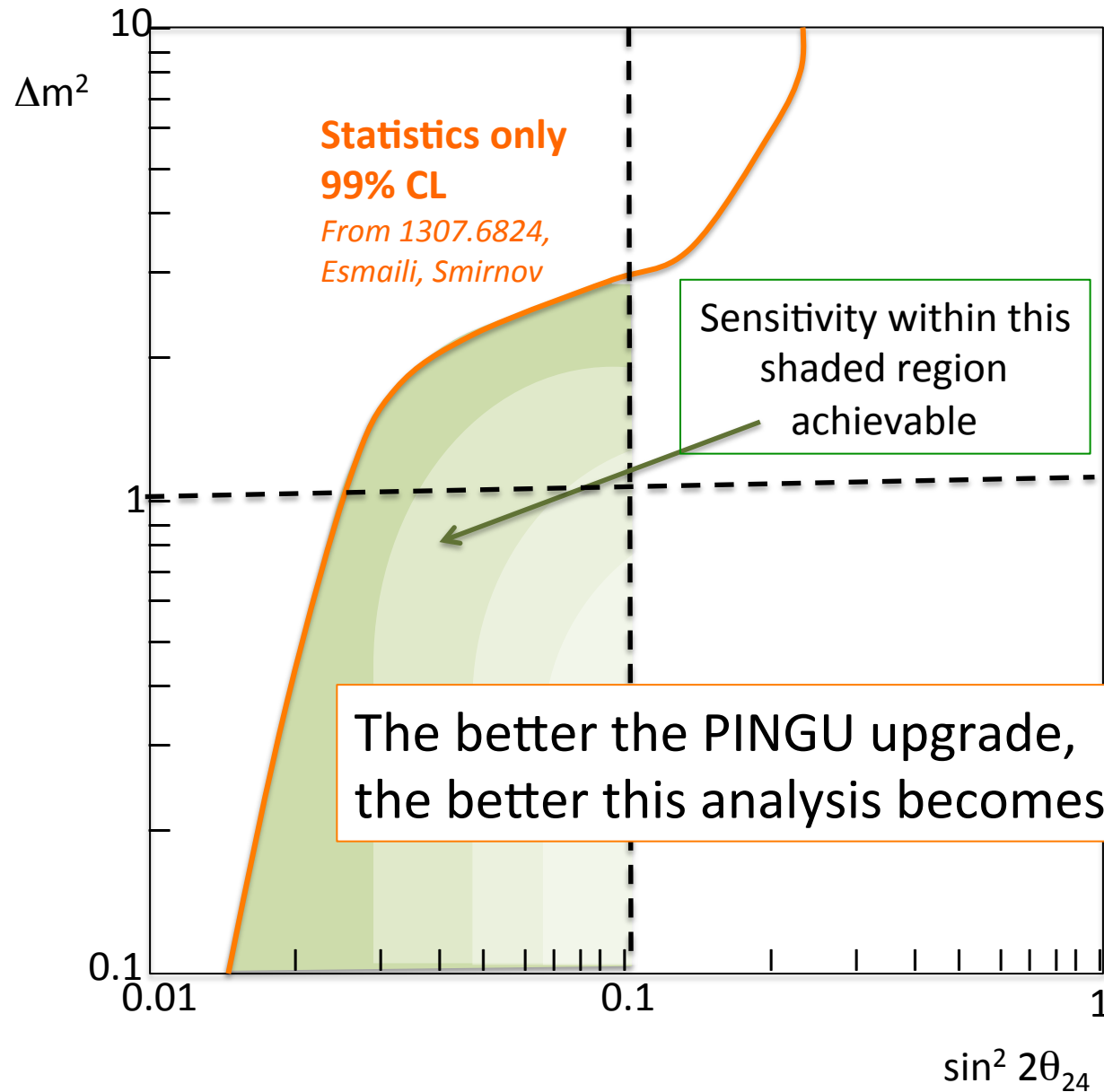


Matter Signal antineutrino, because of the known hierarchy  
→ Access to the antineutrino  $\nu_\mu$  oscillation parameters!

## Three ways this signal is special

1. Wow, that signal is BIG!
2. That signal assumes our QM model of oscillations,  
so this tells us a lot about our sterile model!  
(& would be the first MSW resonance observed!)
3. Any measurement of the osc parameters extracted,  
will be the first measurements in muon antineutrino mode!

Potentially powerful!



First results are already  
on their way!



(Marginalize over  
other parameters)



More Ideas  
can be  
Brought  
to the Table!



What about  $\nu_\tau$  appearance?

→ observable in IceCube/GenTwo, how else?

What about NC disappearance,  
even with more complex beams?

→ e.g. coherent scattering experiments  
at  $\pi/\mu$  DAR sources  
(maybe ultra-high stats  
&  $\nu_e$  disappearance data  
can allow  $\bar{\nu}_\mu$  signal to be  
separated from  $\nu_e$  ...?)

# Conclusion on $\nu_\mu$ disappearance

- We need a strong  $\nu_\mu$  disappearance program.
- There is lots of room for your ideas and “new approaches”!  
Could be room for new  $< \$10\text{M}$  experiments.
- This area of study is not necessarily technically limited.  
We can do a lot without a lot of R&D!

*There is a lot of scientific potential in the sterile neutrino program!*





Let's discuss!